



# NMDB Meeting 2025: Cosmic Ray studies with Neutron Detectors

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### miniTRASGO: Toward a Global Network of Compact RPC Telescopes for Cosmic Ray and Space Weather Monitoring

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# 1. Introduction

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- **Cosmic Rays and Motivation**

- Cosmic rays are high-energy particles (mostly protons and atomic nuclei) arriving from outside the Solar System, producing secondary particles when interacting with Earth's atmosphere (Blasi, 2013; Grieder, 2001).
- Their flux depends on geomagnetic cutoff rigidity, atmospheric absorption, and solar modulation (Dorman, 2004).
- Ground-based detectors monitor cosmic rays to study their temporal and spatial variations, especially important during solar-driven events (Forbush, 1937, 1938).

- **miniTRASGO Project Overview**

- miniTRASGO is a portable, cost-effective cosmic ray detector based on Resistive Plate Chambers (RPCs) (Soneira et al., 2025; in review for publication).
- It aims to complement existing networks by providing directional measurements of charged secondary cosmic rays.
- First stations deployed in Madrid (Spain), Warsaw (Poland), and Mexico, enabling multi-site observations.

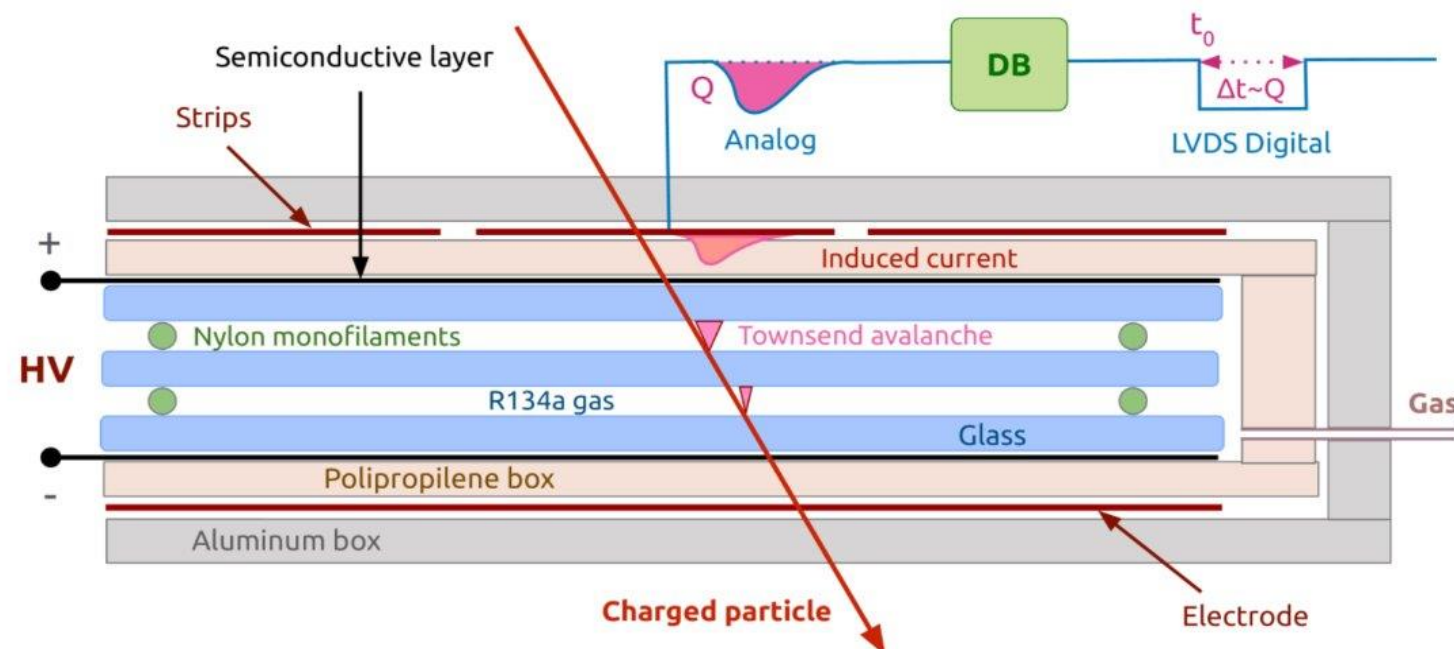
## 2. Materials & Methods (I)

### • Detector Structure

- Four RPC planes arranged vertically within a  $\sim 50$  cm cubic frame.
- Each plane:  $30 \times 30$  cm<sup>2</sup> active area, two-gap configuration enclosed in plastic with high-voltage and gas feedthroughs.
- Operates with pure R134a gas at around 5.5 kV per gap, adjusted for local pressure and temperature (Soneira et al., 2025; in review for publication).

### • Readout System

- Each RPC plane has four copper readout strips (one wide, three narrow) on top plus a reference electrode below.
- Signals from both ends of each strip provide position and timing (Belver et al., 2010).
- Time-to-Digital Converter (TRB3sc) records leading and trailing edges for precise timing and charge measurement.



## 2. Materials & Methods (II)

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- **Signal Processing**

- Front-End Electronics digitize signals above the discriminator threshold, converting them into Low Voltage Differential Signals.
- Charge is inferred via Time-over-Threshold, and positional information comes from the time difference between strip ends.

- **Calibration**

- **Time Offsets:** A custom-made software tool aligns reference times among strips and planes to the few-hundred-ps level.
- **Position:**
  - **Y-coordinate** (perpendicular to strips) determined by which strip(s) fired.
  - **X-coordinate** (along strips) from time difference between front/back ends.
- **Efficiency:** Calculated by checking whether a plane fired when its three companion planes detected a crossing particle, yielding ~95% efficiency (Soneira et al., 2025; in review for publication).

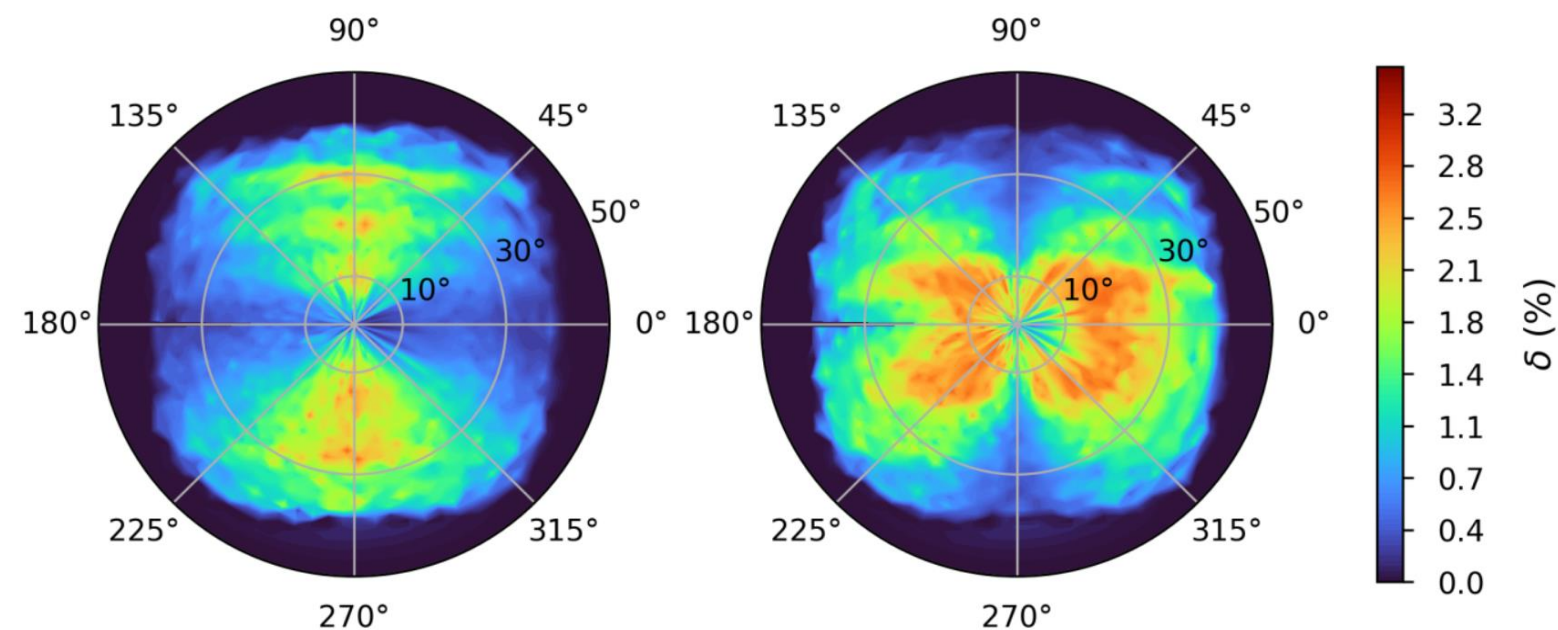
### 3. Results (I)

- **Angular and Charge Measurements**

- Angular resolution better than 3% for muons above  $\sim 1$  GeV, where multiple scattering is minimal.
- Charge distributions show single-strip signals dominate ( $\sim 90\%$ ), with multi-strip “clusters” mainly due to adjacent strip sharing or crosstalk.

- **Atmospheric Corrections**

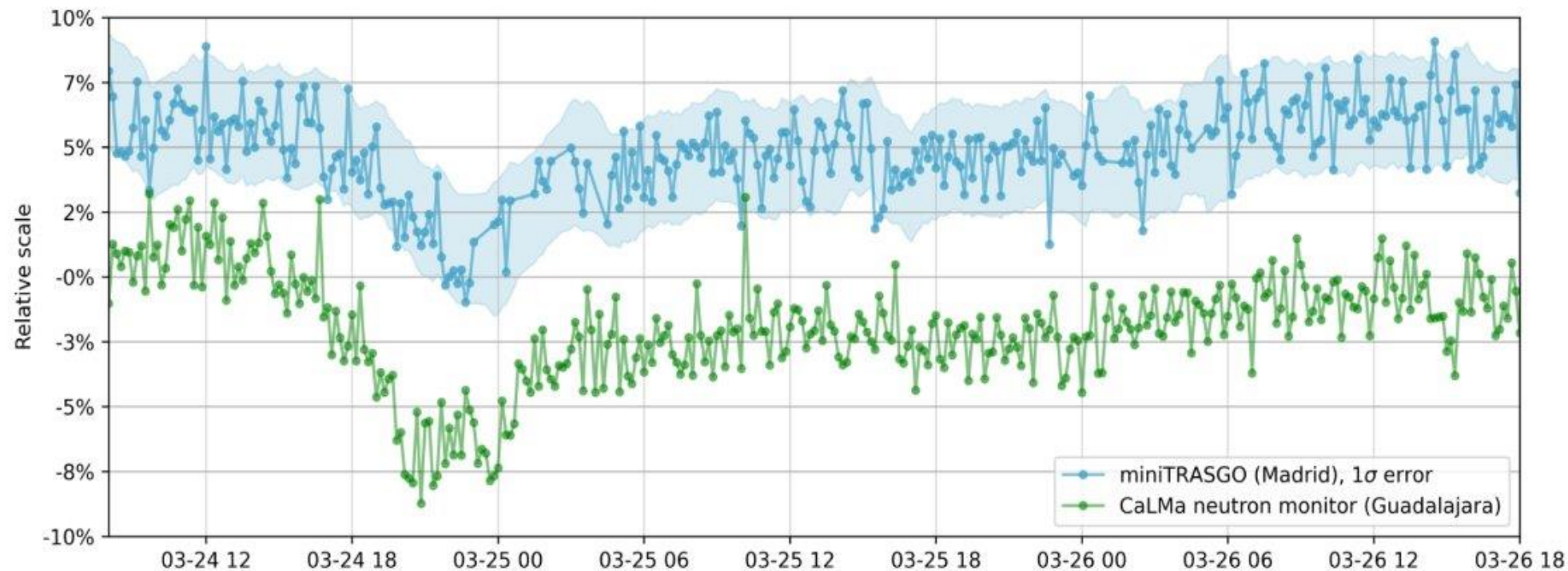
- Pressure changes significantly affect rates. A barometric coefficient of about  $-0.215\%/mbar$  (Madrid station) is applied to correct raw count rates.



### 3. Results (II)

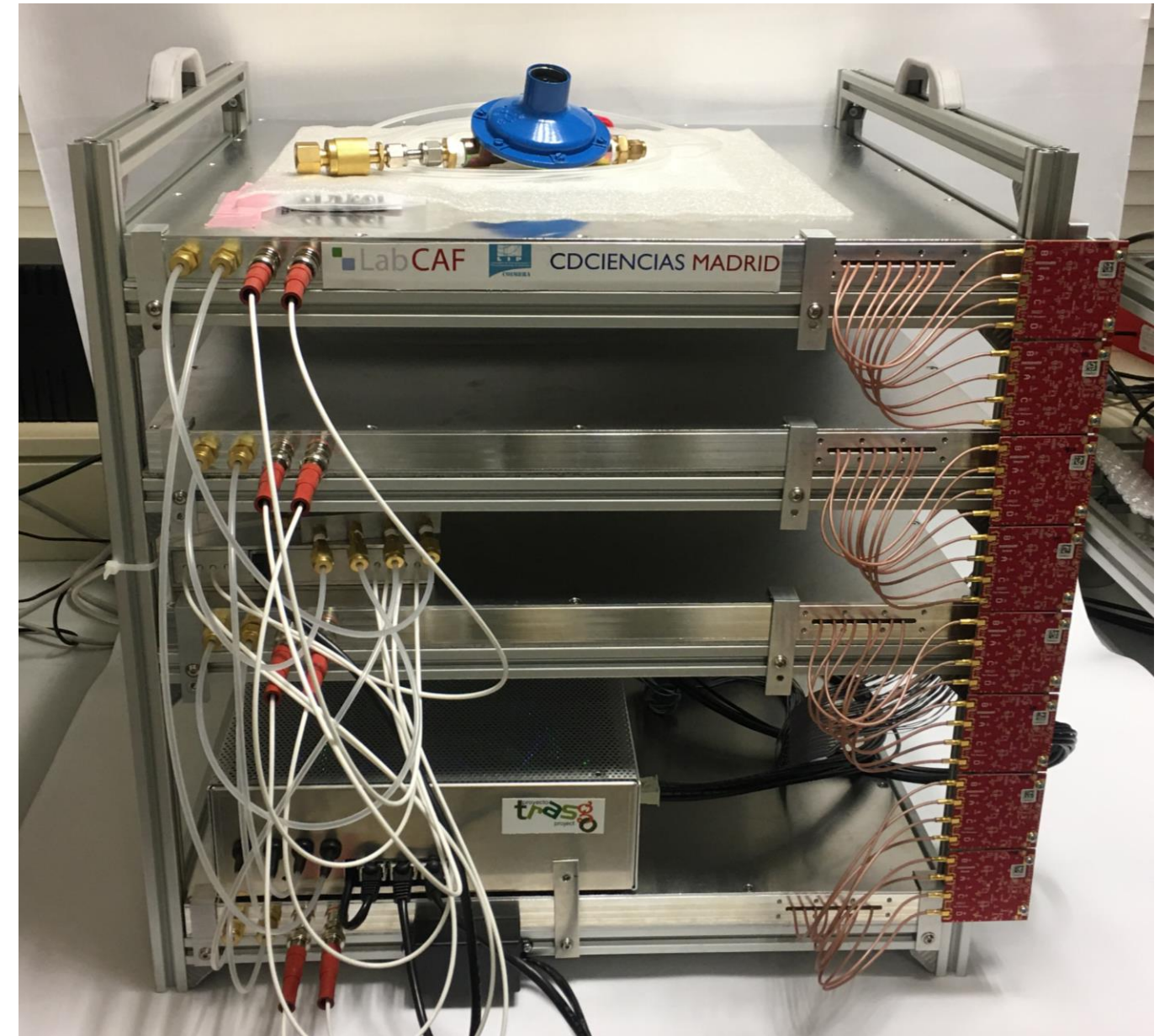
- **Forbush Decrease Detection**

- miniTRASGO observed a  $\sim 5\%$  drop in cosmic ray intensity during the March 2024 Forbush Decrease, consistent with CaLMA neutron monitor data in Spain (Steigies et al., 2008).
- Demonstrates miniTRASGO capability to track transient cosmic ray events for space weather studies.



## 4. Conclusions

- **Performance**
  - Compact and easily deployable system with good angular ( $<3\%$ ) and time resolution ( $\sim 300$  ps).
  - Consistent efficiency ( $\sim 95\%$ ) under varying environmental conditions.
- **Implications for Global Monitoring**
  - Complements existing neutron monitor and muon detector networks by adding directional sensitivity and broader geographic coverage.
  - Ongoing and future deployments will enable multi-site comparative data on cosmic ray flux variations and solar-terrestrial interactions (Soneira et al., 2025; in review for publication).



## 5. References

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