



# Gammapy: current status and future plans

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on behalf of the Gammapy dev team

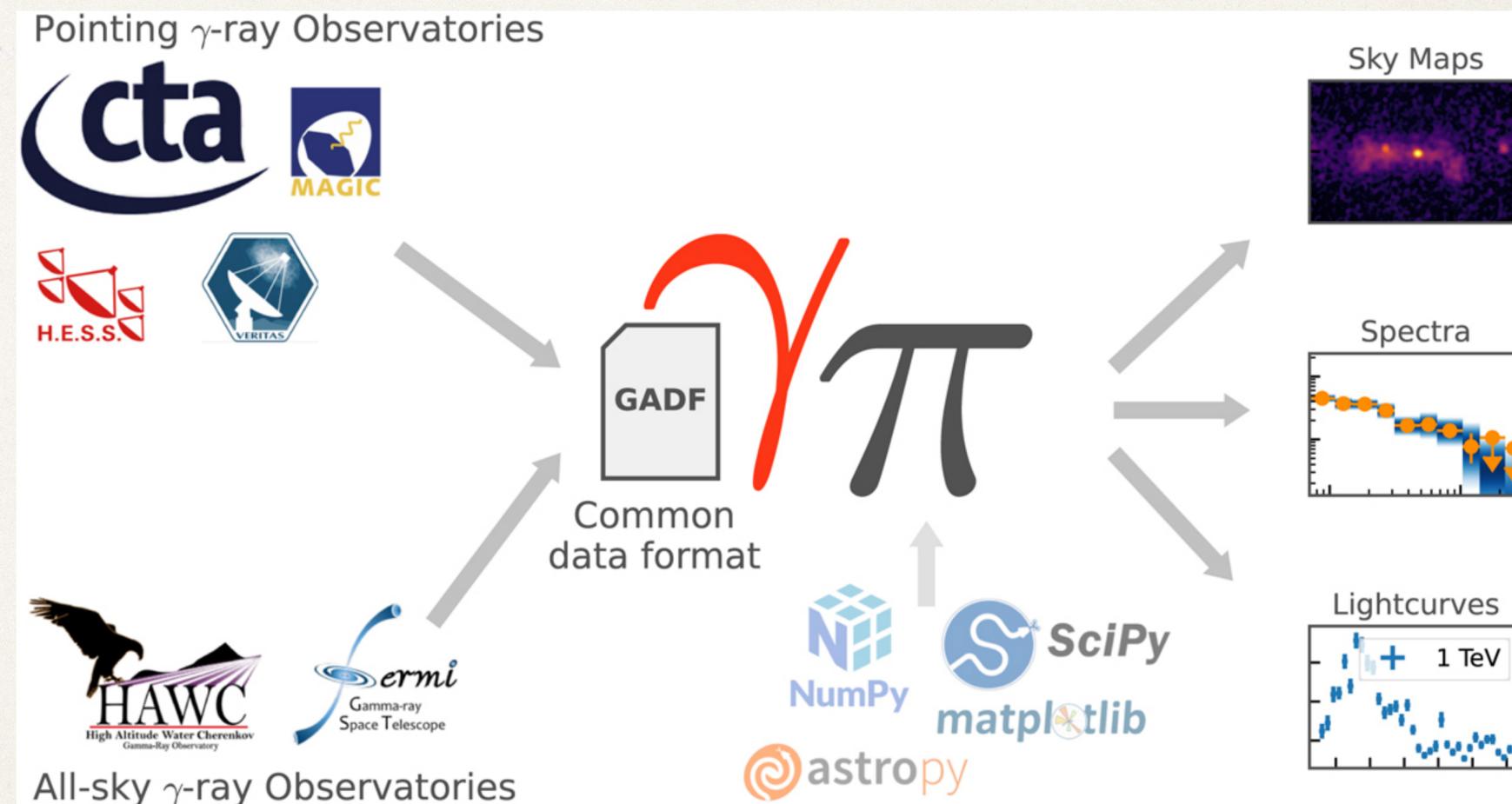
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# Gammapy: high level analysis tool

- ❖ 2013: Started as an open source python package for Gamma-ray astronomy
- ❖ 2021: Library for the CTAO science tools
  - ❖ Open IACT observatory
- ❖ Adopted by H.E.S.S., MAGIC, VERITAS, HAWC, LST1, ASTRI, SWGO
- ❖ Supports Fermi-LAT analysis
- ❖ Data to be serialised according to the GADF/VODF formats



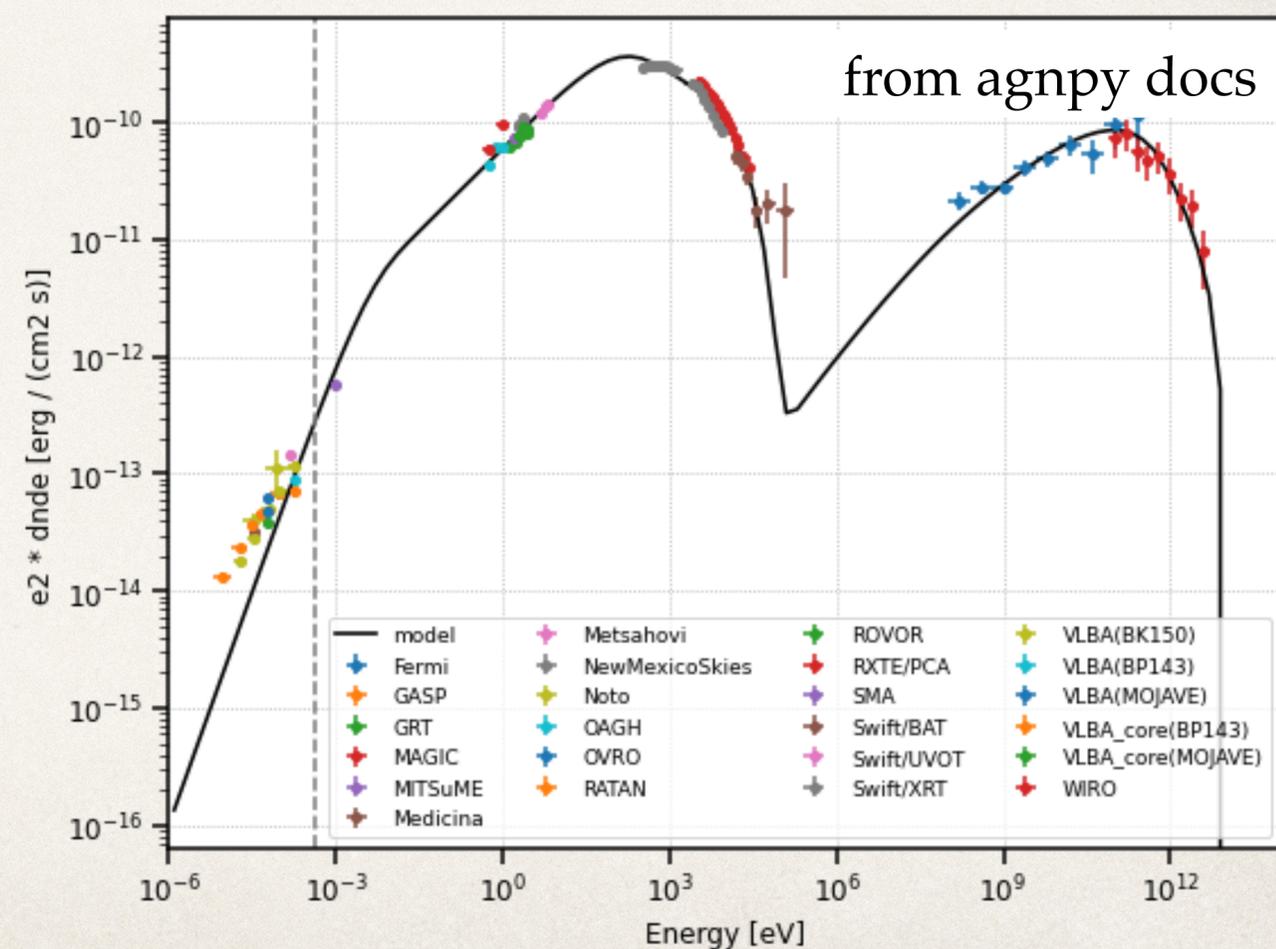
# Functionalities

- ❖ Data handling from multiple instruments
- ❖ IRF access and manipulation
- ❖ Background estimation (ring, reflected, FoV...)
- ❖ Joint likelihood from same or different instruments
- ❖ Imaging (flux, ts maps, ...)
- ❖ Spectral analysis
- ❖ Temporal analysis
- ❖ Parameter estimation
- ❖ Observation simulations
- ❖ High level interface
- ❖ Common catalogs



# Common multi-wavelength approach

- ❖ Extract flux points from individual instruments
  - ❖ Different frameworks
  - ❖ Usually considering simplistic models (eg: Power law)
- ❖ Fit (or plot) physical models on top
  - ❖ Systematics (if any) added in quadrature
  - ❖ Gaussian errors assumed



# Solution: Fit directly on the counts

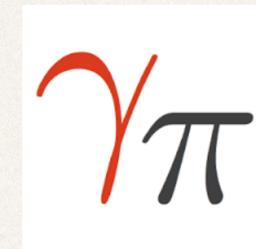
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Forward folding directly with emission models



3ML: Wrappers around each native tool to obtain the likelihood functionality.

- ✓ Uses native tools
- All tools need to 'coexist' in the same environment

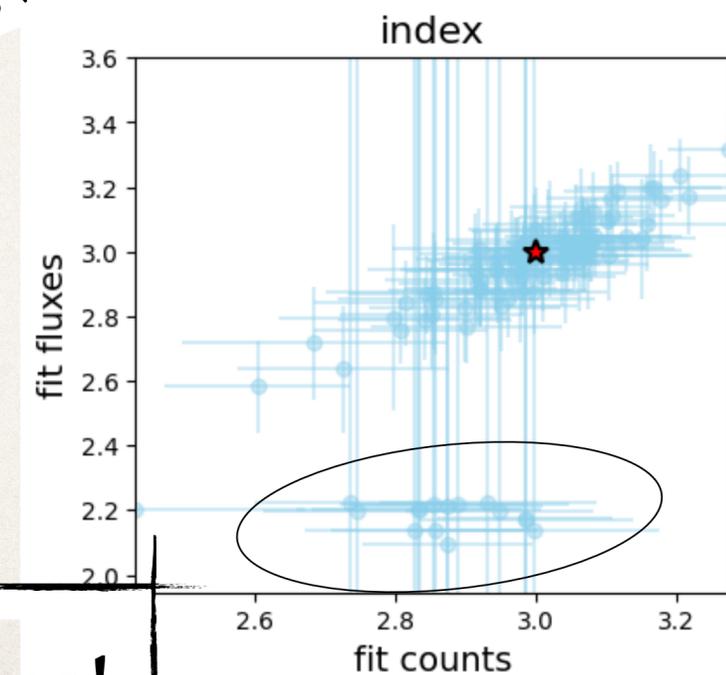
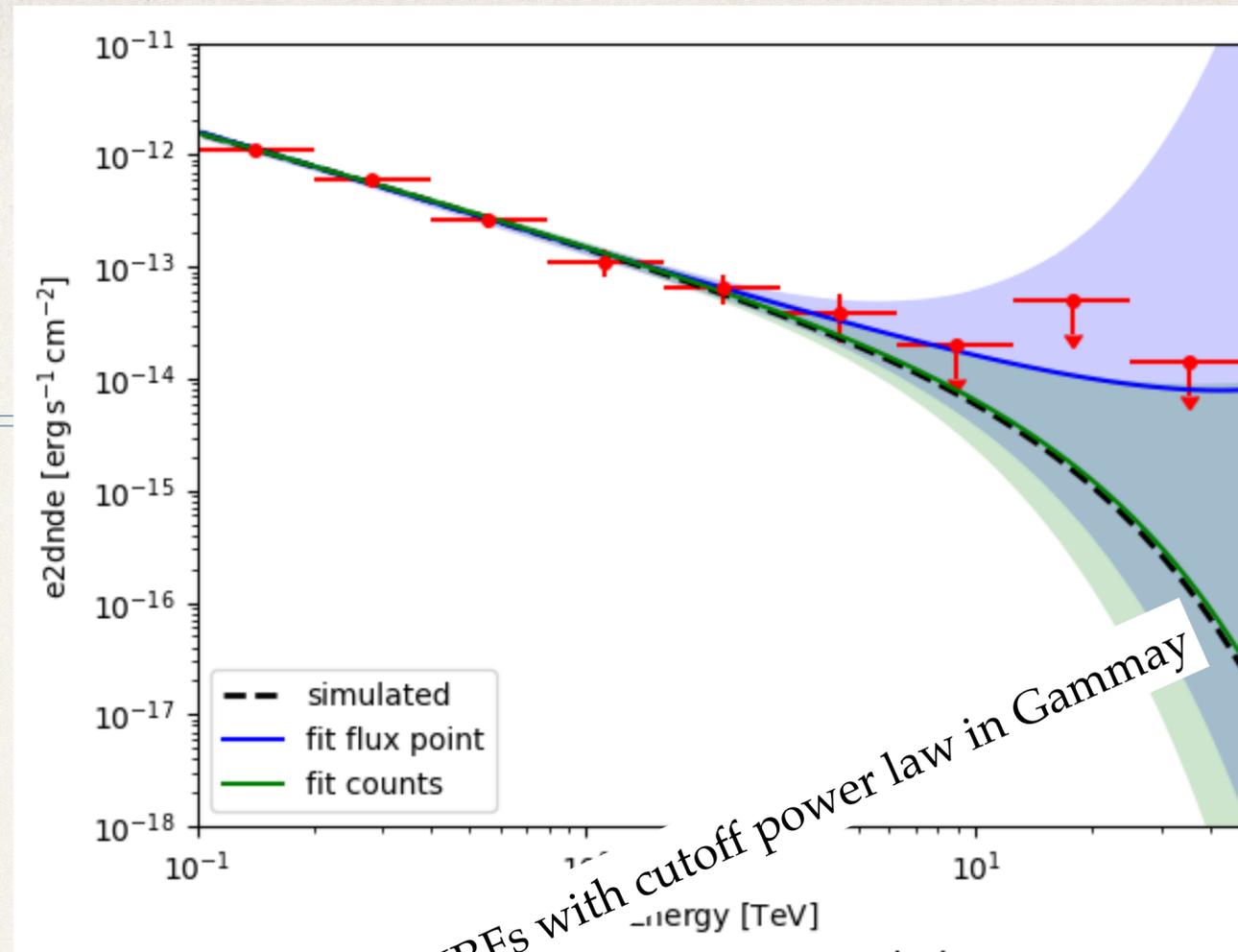


Gammapy: Convert the data and instrument description to a common format.

- ✓ Common likelihood code
- ✓ Handling of inter-instrument effects
- Need to convert the data and the instrument responses into common format

# Not ideal

1. Fitting flux points may be biased
  - Lose underlying features depending on extraction type
2. Unclear Handling of upper limits
  - Upper limits depend on underlying model assumption
3. X-rays: xspec is NOT designed to produce flux points - "hacks" used
4. Ignores flux points correlations
  - Specially important if Flux points were obtained through unfolding
5. Statistical assumption: Gaussian statistics

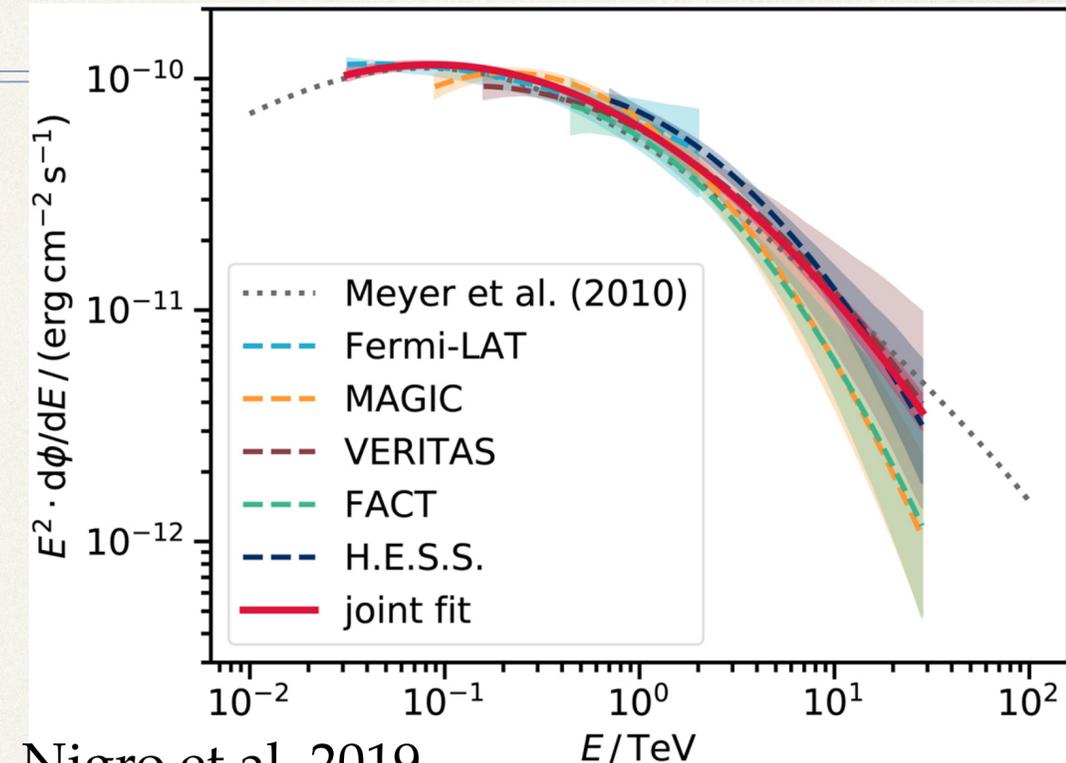


Reconstructed parameters may differ from true value!

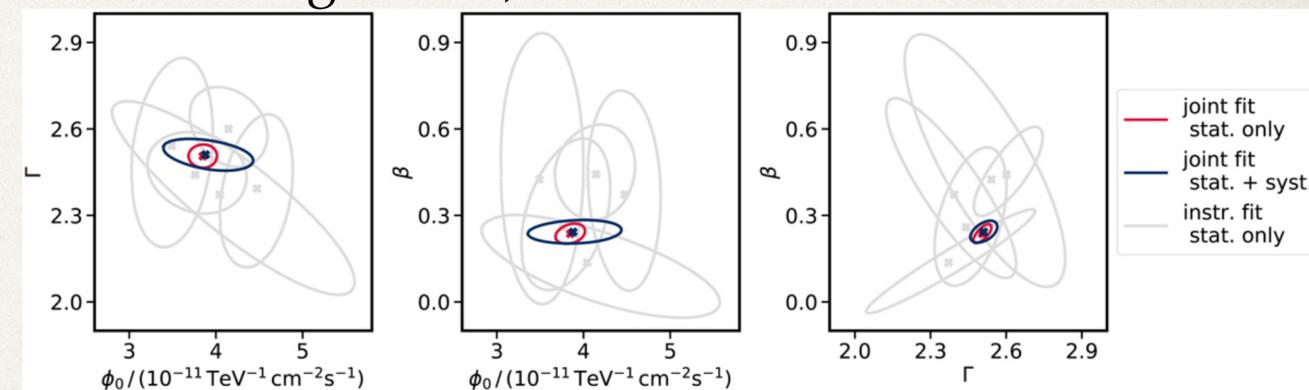
# Joint likelihood for Gamma-ray instruments

- Simultaneous fitting from various instruments at the counts level
  - ✓ Different types of **models** and **likelihoods** for **different datasets**, eg
    - ❖ Fermi-LAT: Multiple sources in the FoV; cash likelihood
    - ❖ IACT: src model+ absorption, on-off background, wstat likelihood
    - ❖ Also allows to use precomputed flux points; chi2 stat
  - ✓ Efficient book-keeping between different models and datasets
  - ✓ Systematics included as nuisance parameters in likelihood

```
FIT_STATISTICS_REGISTRY = {  
    "cash": CashFitStatistic,  
    "wstat": WStatFitStatistic,  
    "chi2": Chi2FitStatistic,  
    "distrib": Chi2AsymmetricErrorFitStatistic,  
    "profile": ProfileFitStatistic,  
    "cash_weighted": WeightedCashFitStatistic,  
}
```



Nigro et al, 2019

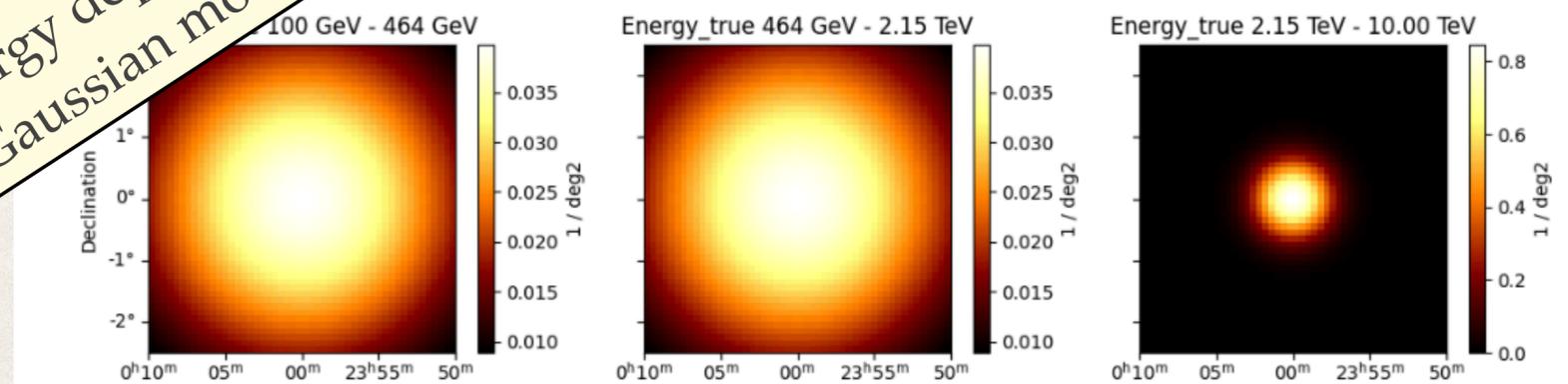
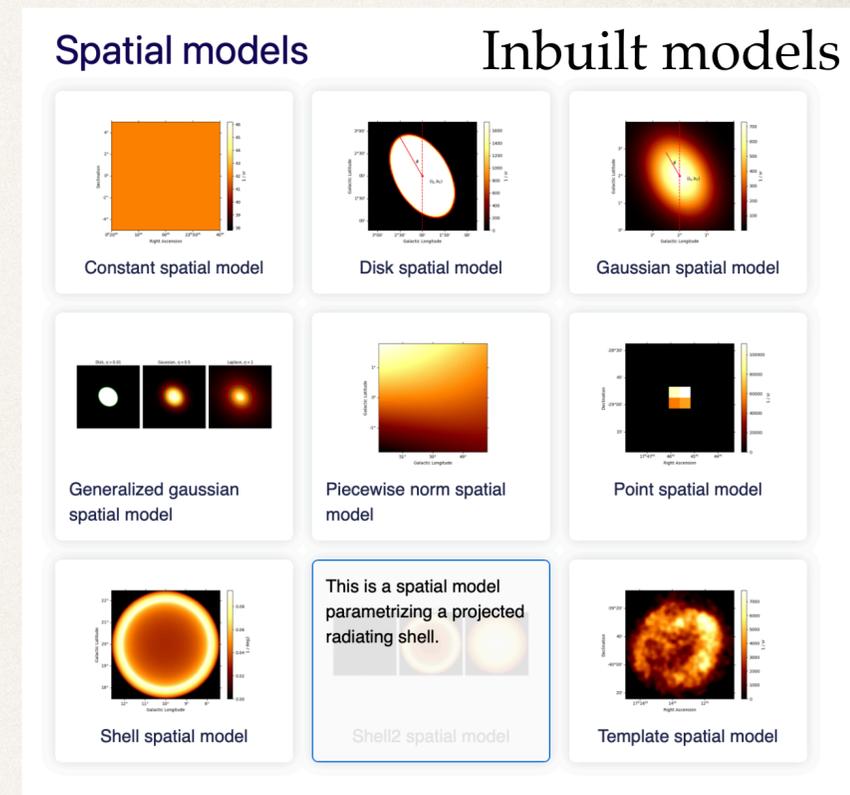


# Support for user contributed models

- Inbuilt support for common analytical model definitions
- Easy to incorporate your own models
  - As a `Custom model`
    - Analytic models
  - As a `TemplateModel`
    - Precomputed n-dim grid + interpolation
    - Useful for computationally heavy models (eg: hadronic models)
- Interoperability with physical models
  - Dark matter models (COSMIXS, PPPC4)
  - Blazar models (JetSET, Agnpy)
  - Hadronic models (Naima, GAMERA)

$$f_{Src} = f_{Spectral}(E) \cdot f_{Spatial}(E, l, b) \cdot f_{Temporal}(t)$$

An analytic  
energy dependent  
Gaussian model



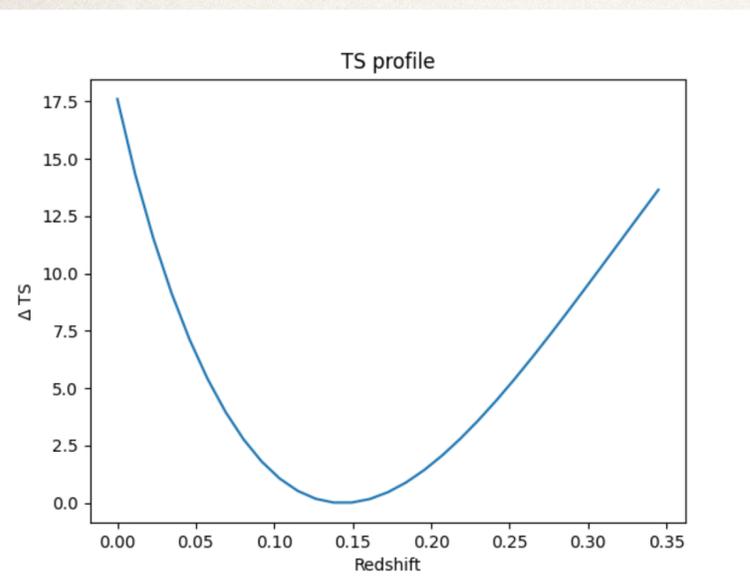
# Inbuilt support for EBL absorption

- ❖ EBL models can be multiplied with spectral models to account for the absorption
- ❖ Redshift and absorption scale are model parameters
  - ❖ You can fit these values directly!

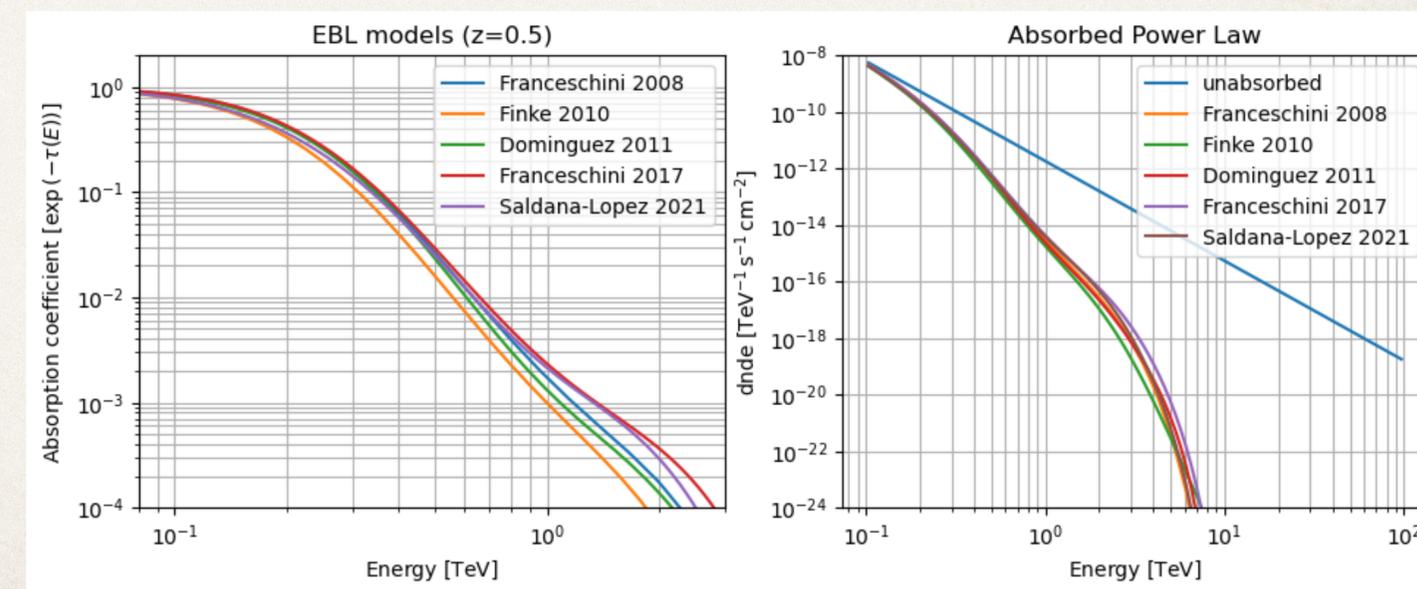
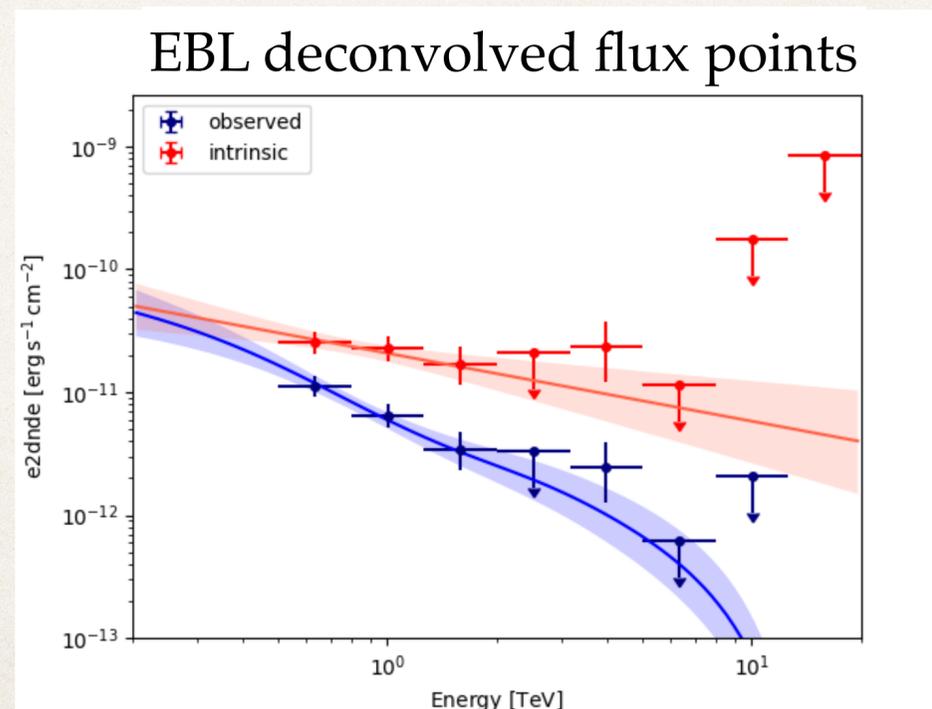
$$\exp(-\alpha \times \tau(E, z))$$

Scalable norm

Predicted optical depth

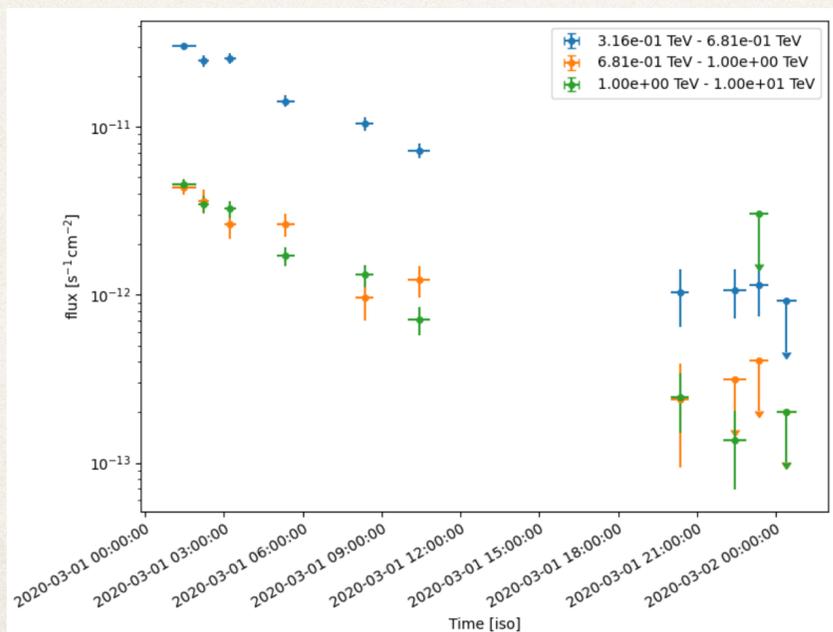


Likelihood profile of redshift

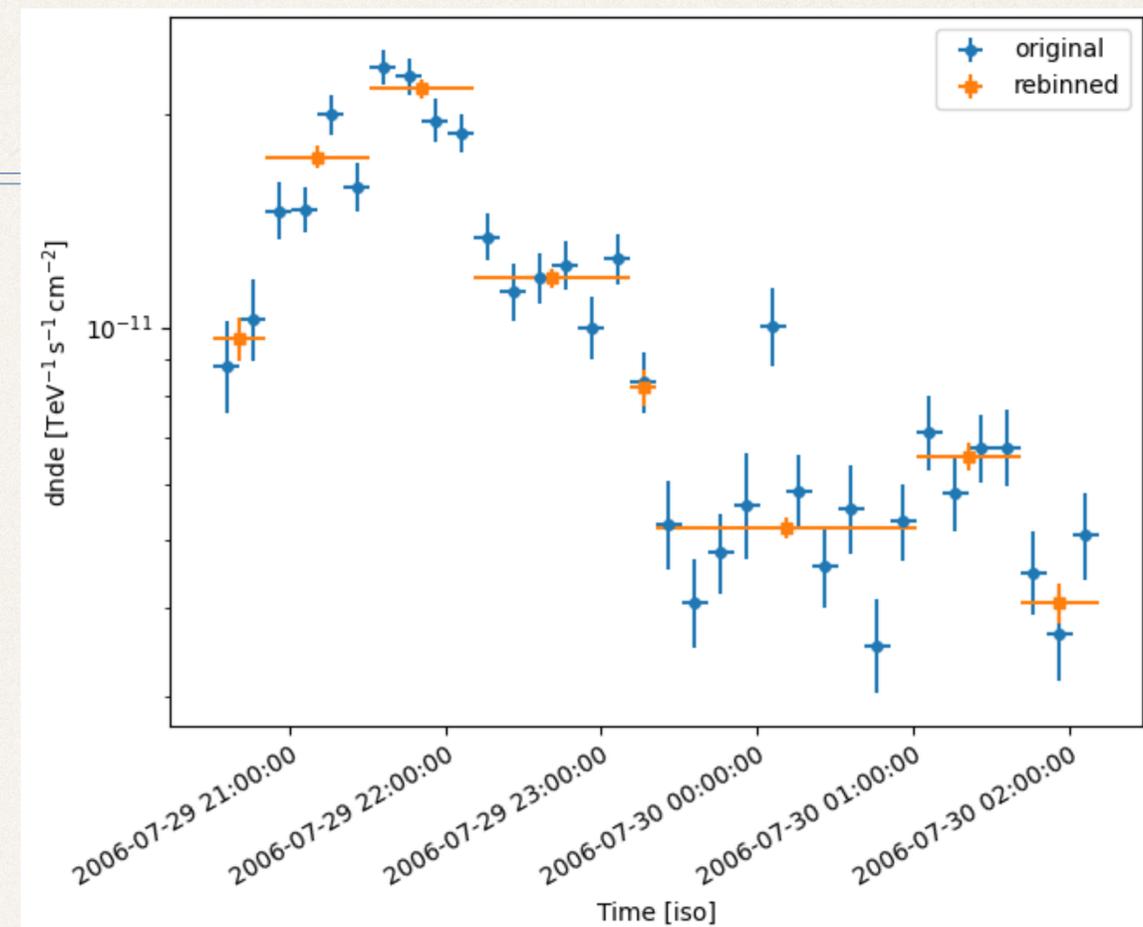


# Temporal analysis

- ❖ Creation of lightcurves
  - ❖ Rebin lightcurves on the fly using likelihood profiles
- ❖ Simulation and fitting of light curves
  - ❖ Timmer & König and Emmanoulopoulos algorithms
  - ❖ Fitting flares with temporal models
- ❖ Inbuilt variability estimation tools



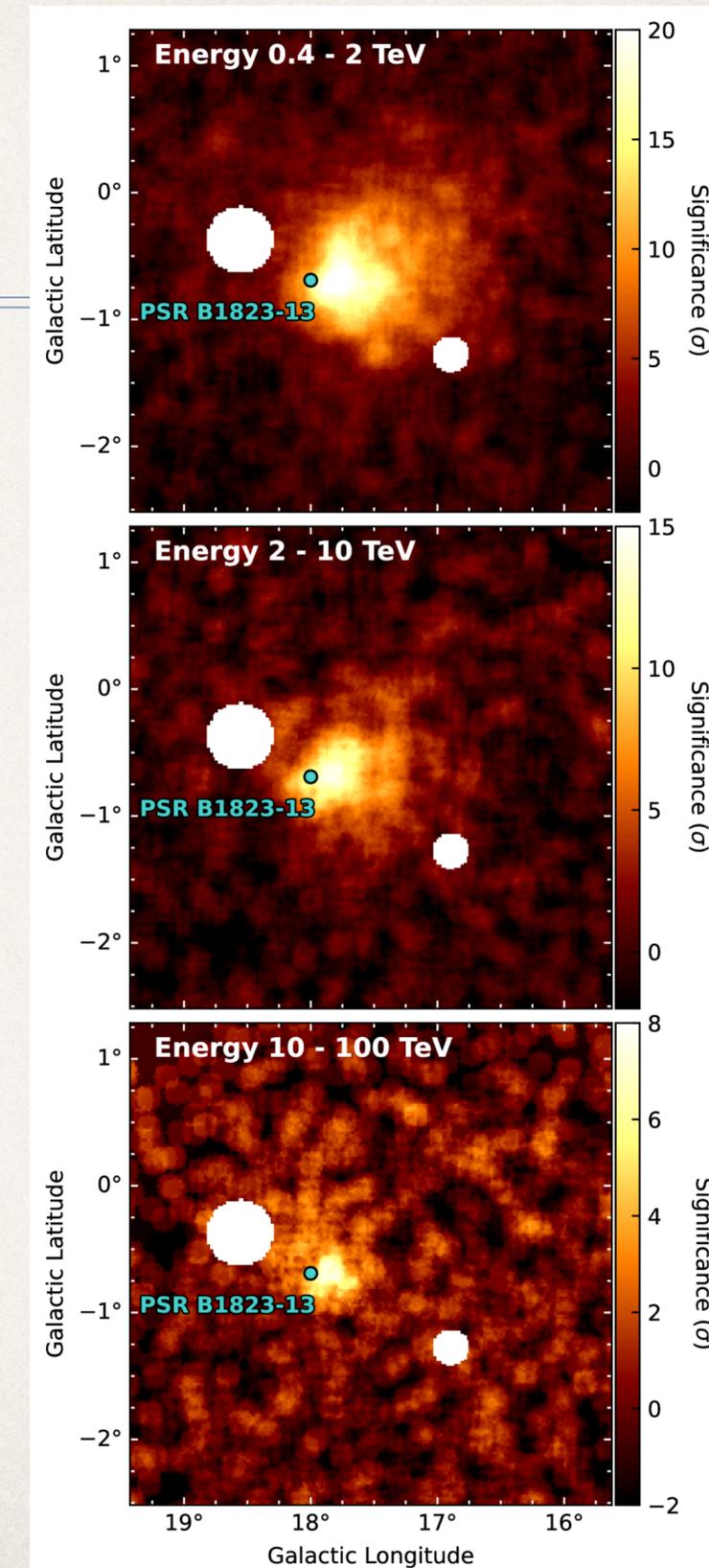
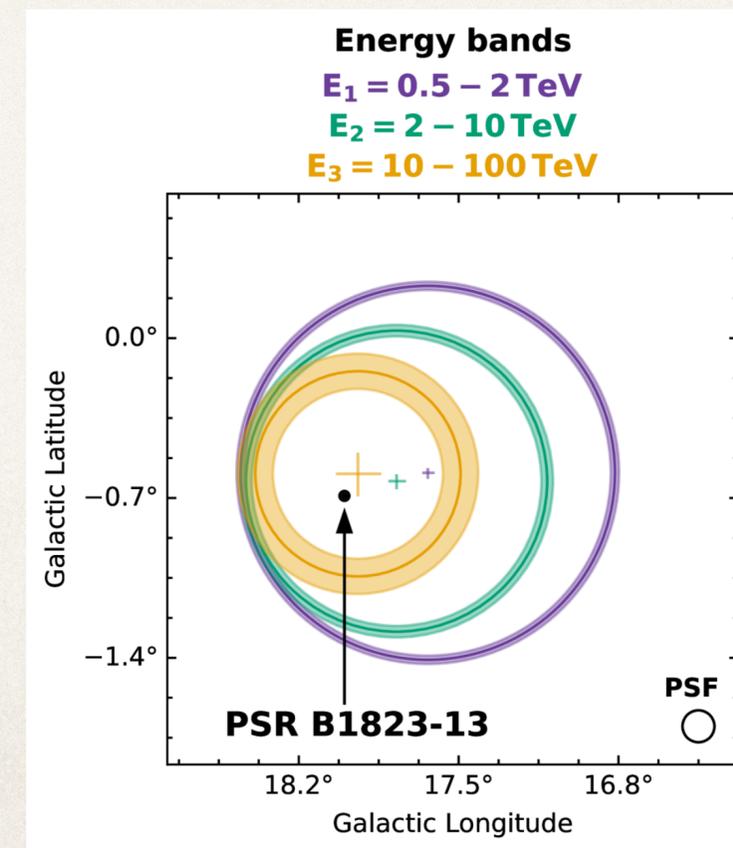
Simultaneous  
  
 Spectro-temporal fitting



name	value	unit	error	frozen
index	2.9558E+00		3.131E-02	False
amplitude	1.0292E-11	cm-2 s-1 TeV-1	3.541E-13	False
reference	1E+00	TeV	0E+00	True
t0	5.8447E+00	h	2.002E-01	False
t_ref	5.8909E+04	d	0E+00	True

# Energy dependent morphology

- ❖ Estimators to compute flux and significance maps in different energies
- ❖ Estimator to quantify if source morphology changes with energy



# Obtaining “best” parameters

❖ Gammapy offers several statistical methods to *estimate* ‘best’ parameters from the data:

❖ **Maximum likelihood estimation (MLE):**

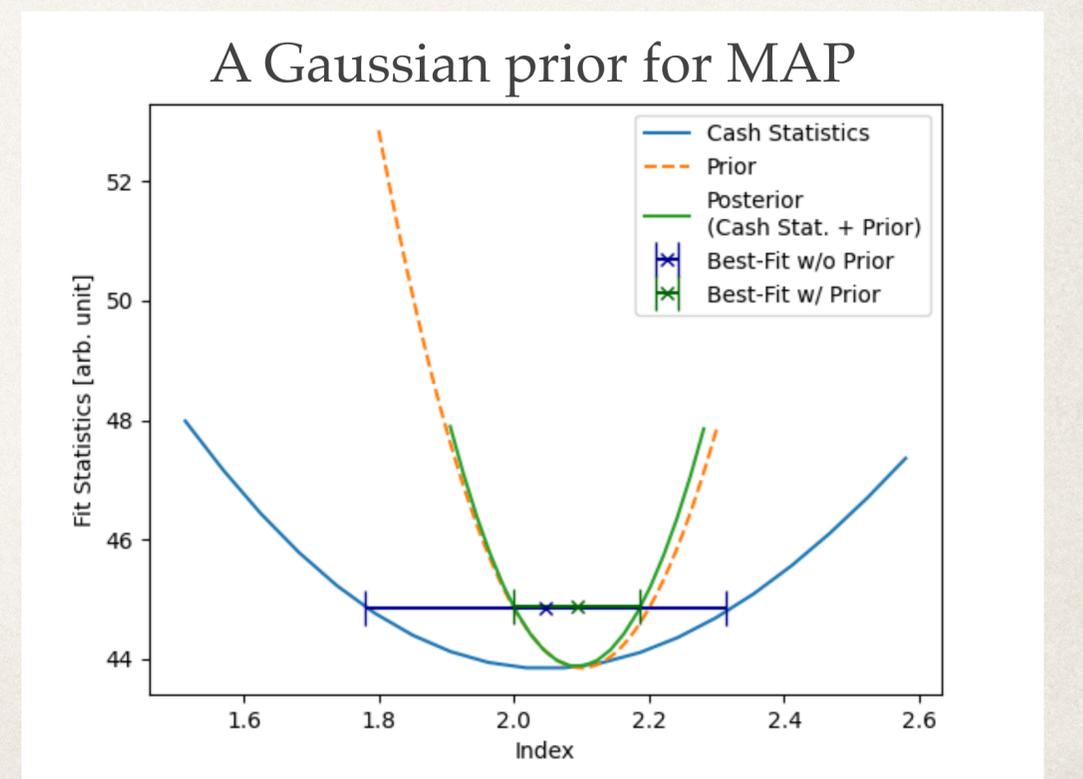
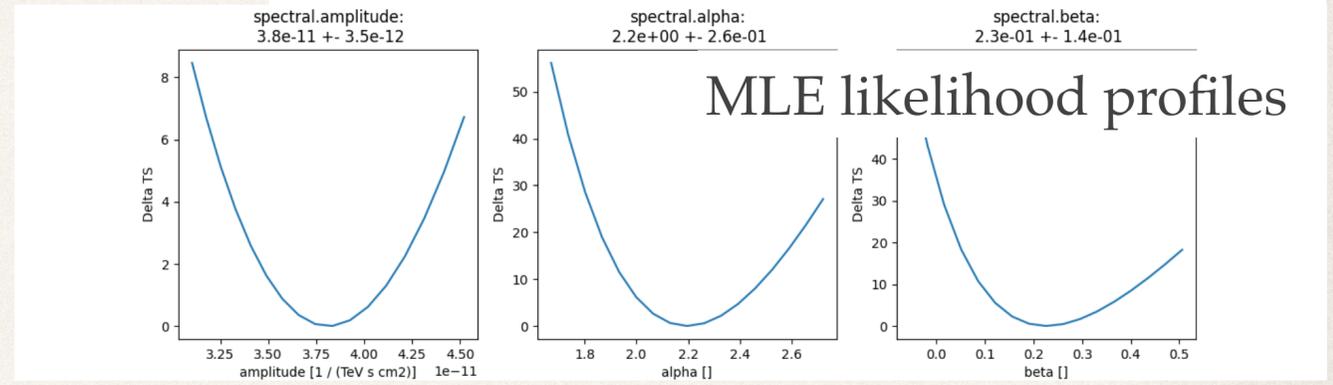
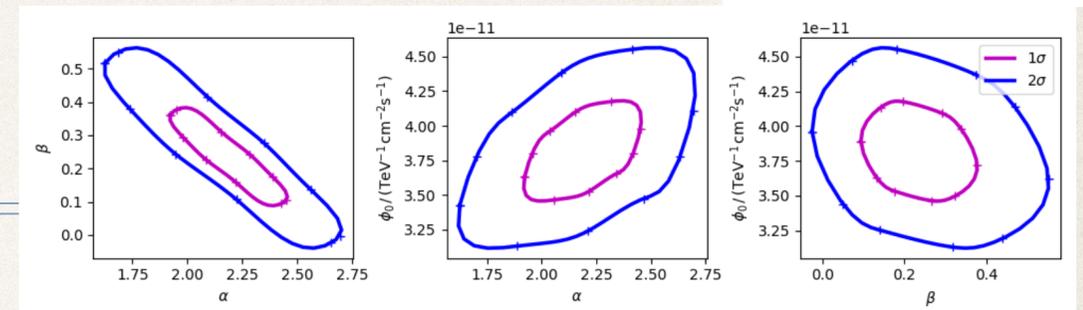
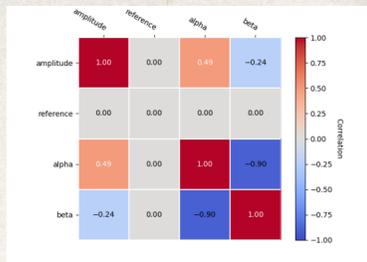
- ❖ Frequentist inference of parameters from sample data
- ❖ Hypothesis testing, Confidence intervals...
- ❖ Prior distributions of parameters are uniform in the region of interest

❖ **Maximum A Posteriori estimation (MAP):**

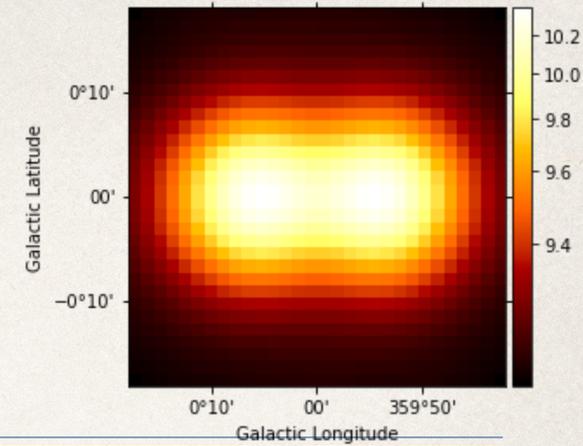
- ❖ Incorporating a *Prior* density over the quantities one wants to estimate
- ❖ Maximise the a posteriori likelihood

❖ **Bayesian inference**

- ❖ Inference of the best a posteriori parameters
- ❖ Not associated to the “best model” in the Frequentist sense, but rather to the most probable given a set of parameters’ priors.
- ❖ Very powerful in case of
  - ❖ non-Gaussian degeneracies
  - ❖ larger number of parameters,
  - ❖ map likelihood landscapes with multiple solutions
- ❖ Can be more computationally intensive

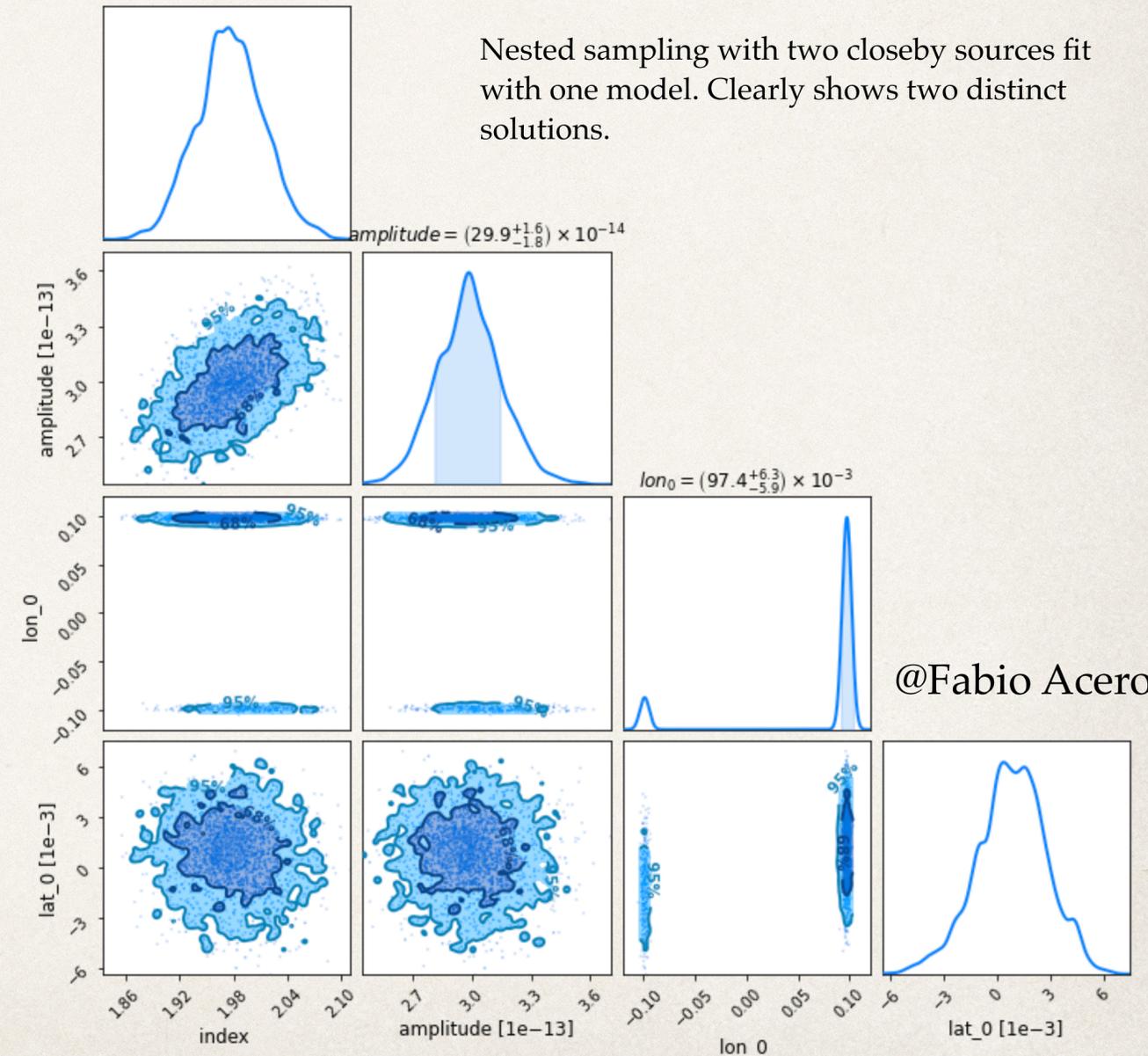


# Bayesian analysis



- Support for Bayesian analysis with Nested Sampling

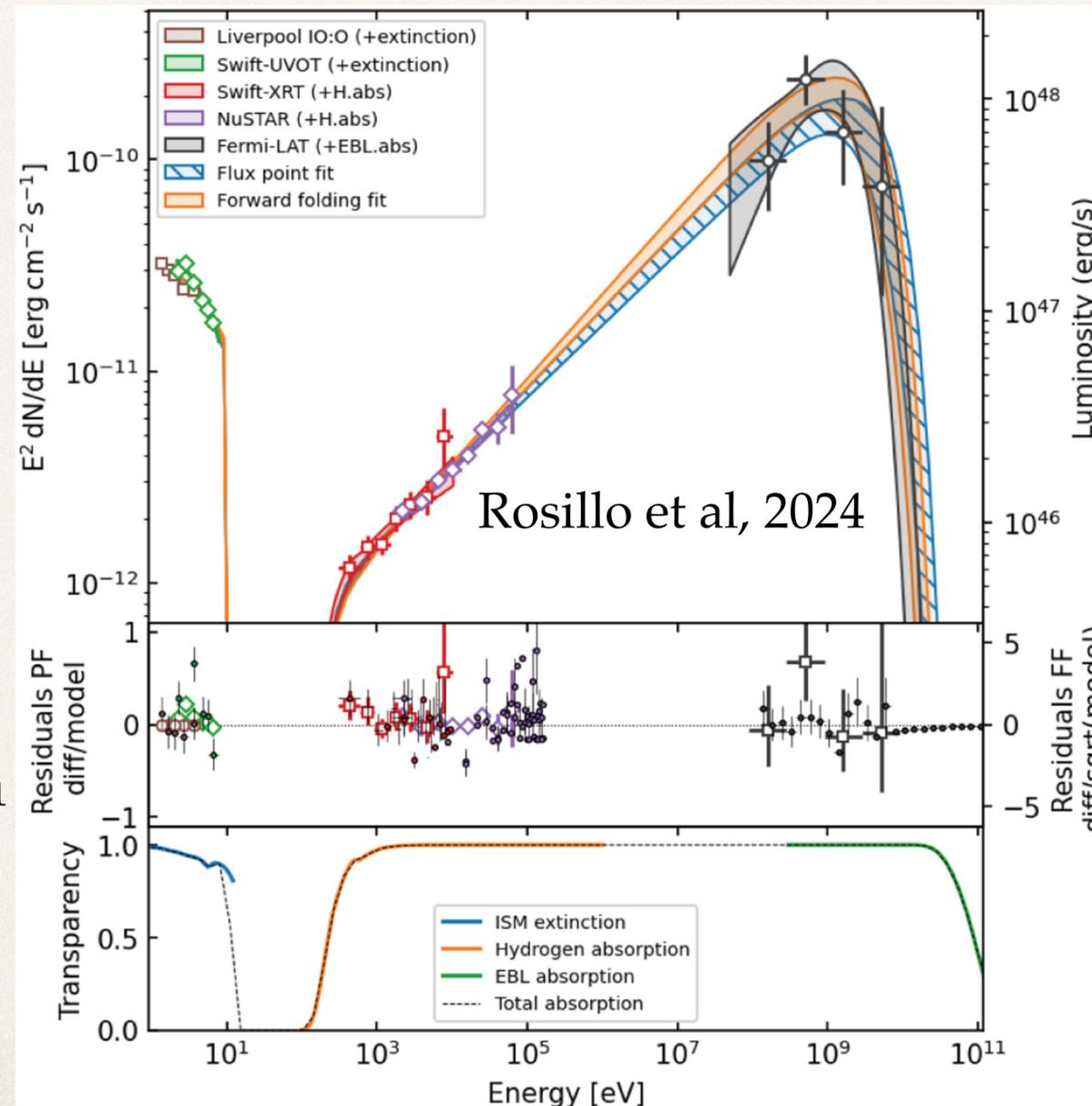
- ✓ No initial points
- ✓ No burn-in period (eg: as in MCMC)
- ✓ No issues with multi-modal distributions
- ✓ Obtain the full posterior distribution
- ✓ Calculates the Bayesian evidence  $Z$ : compare non-nested models
- Need to define priors
- Slower



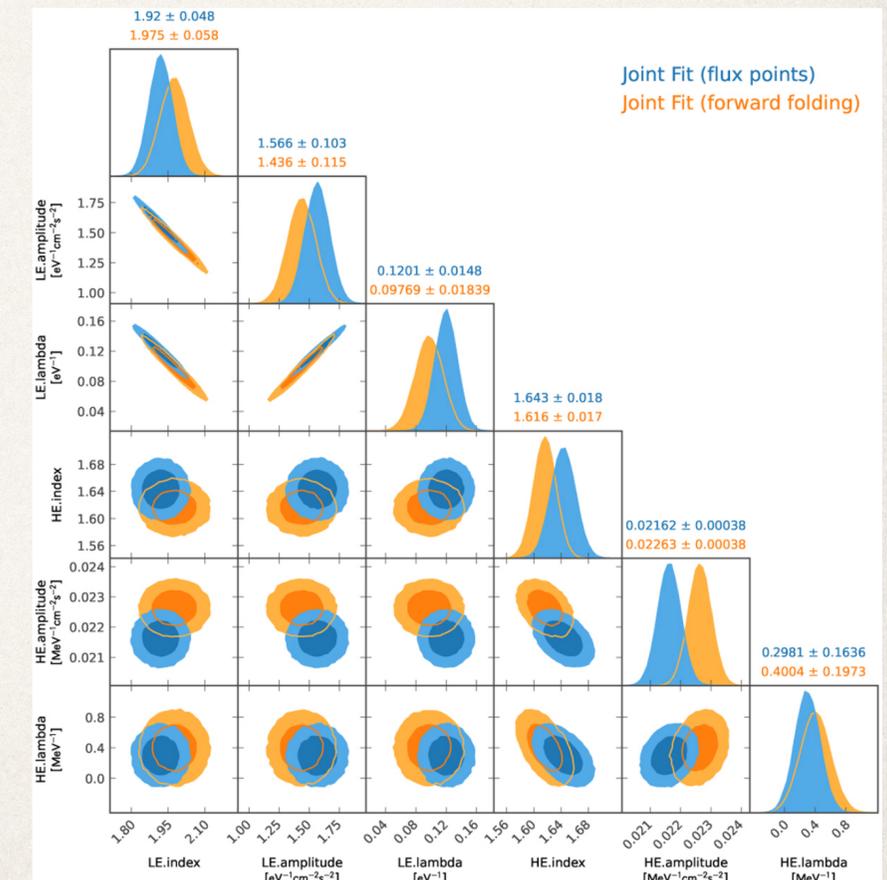
# Beyond gamma-ray analysis: multiwavelength

<https://github.com/gammapy/gammapy-mwl>

- ❖ Extension to X-rays and Optical/UV
- ❖ Multi-instrument, multi-wavelength forward folding fit for OP 313 over 10 decades in energy
  - ❖ Multiple sources in LAT + diffuse components
  - ❖ Instrumental background in NuSTAR
  - ❖ Absorption components: EBL, hydrogen, reddening

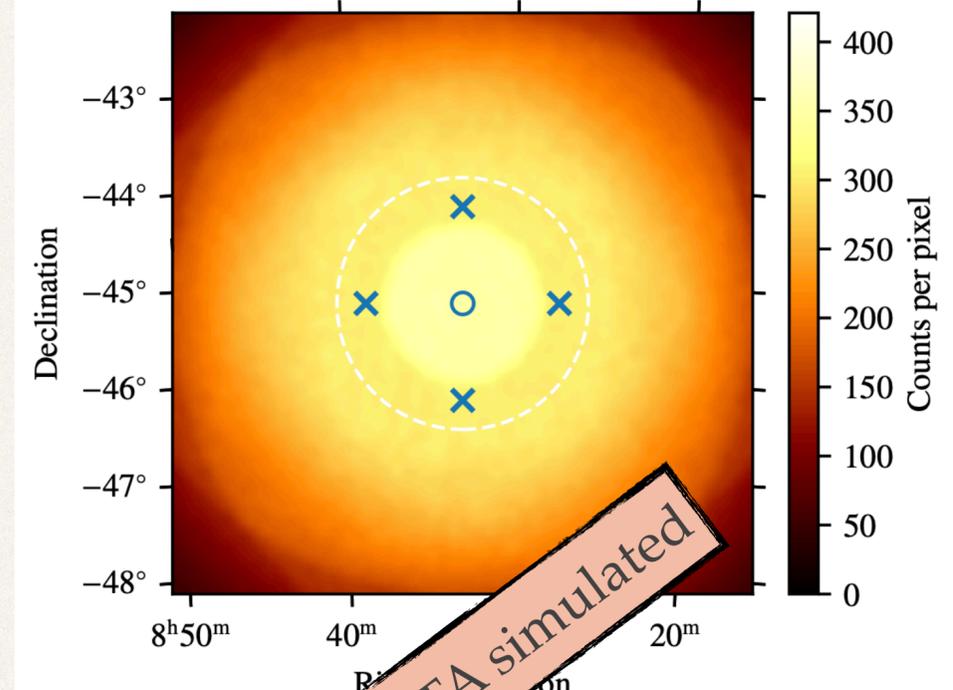
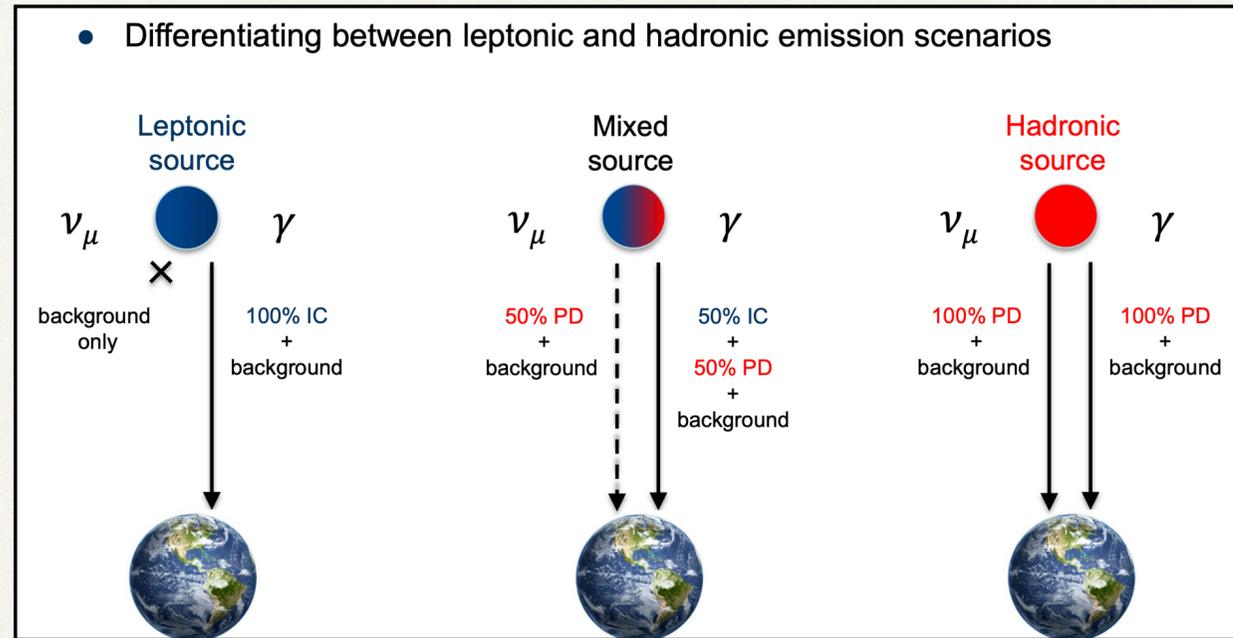
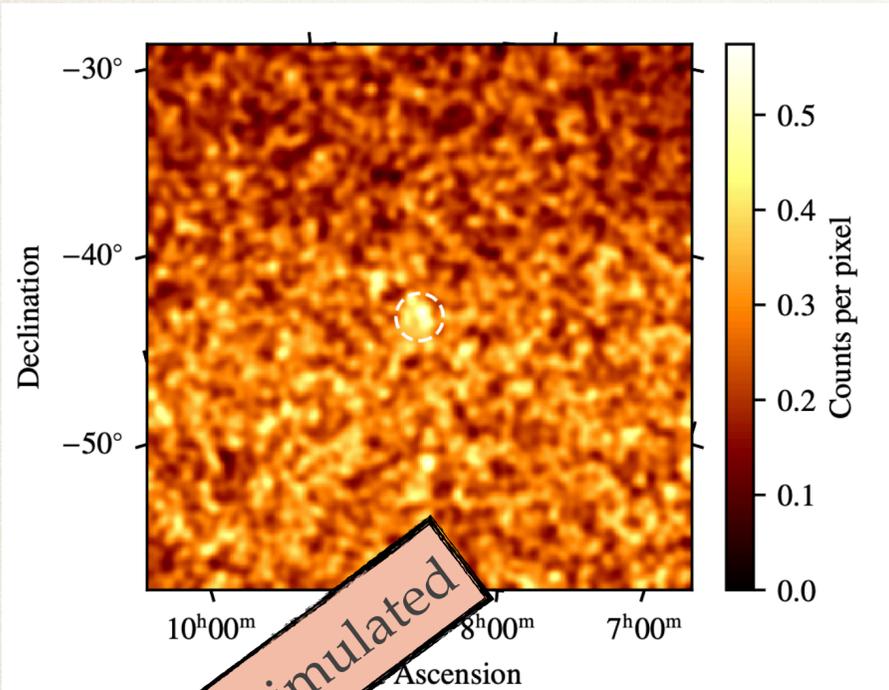


Corner plot from cov matrix

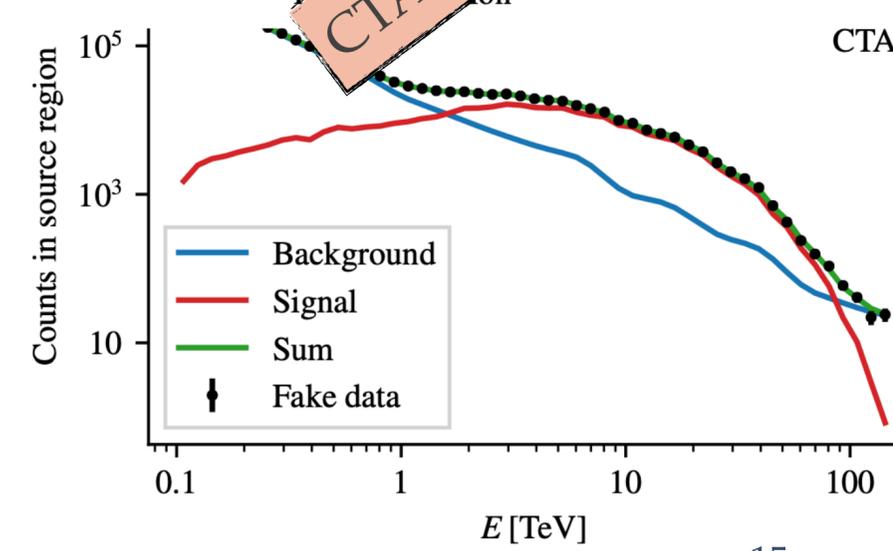
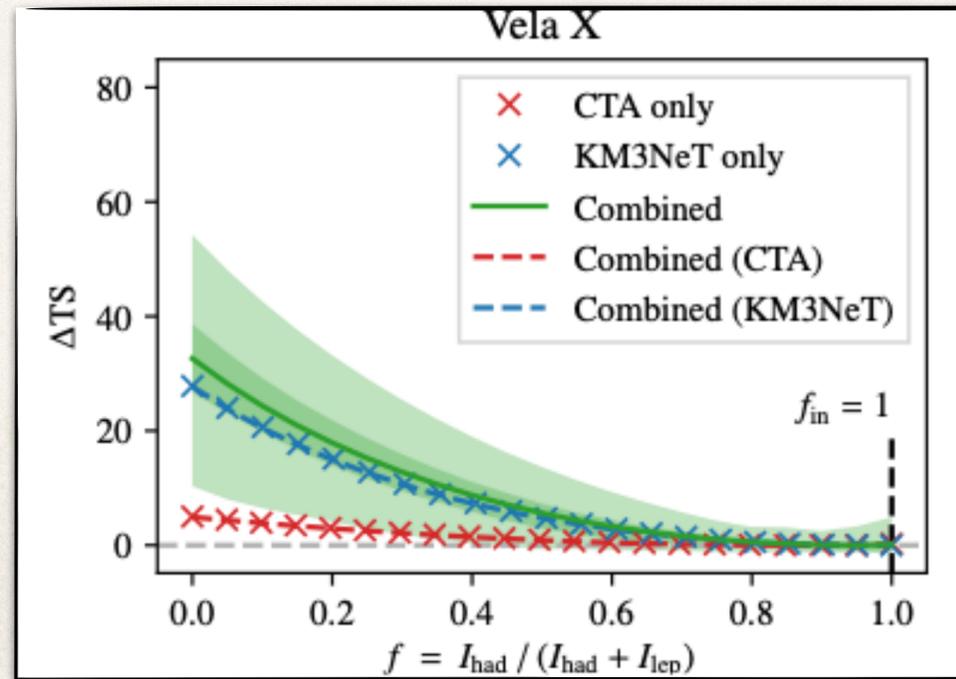
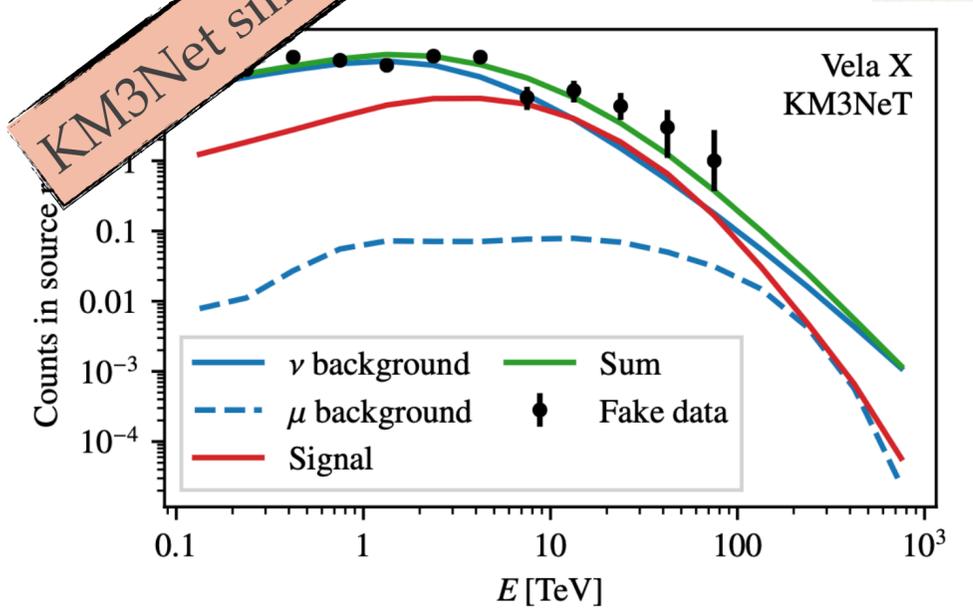


# Beyond EM: Multimessenger

## Prototype joint KM3NET and CTAO analysis



KM3Net simulated



CTA simulated

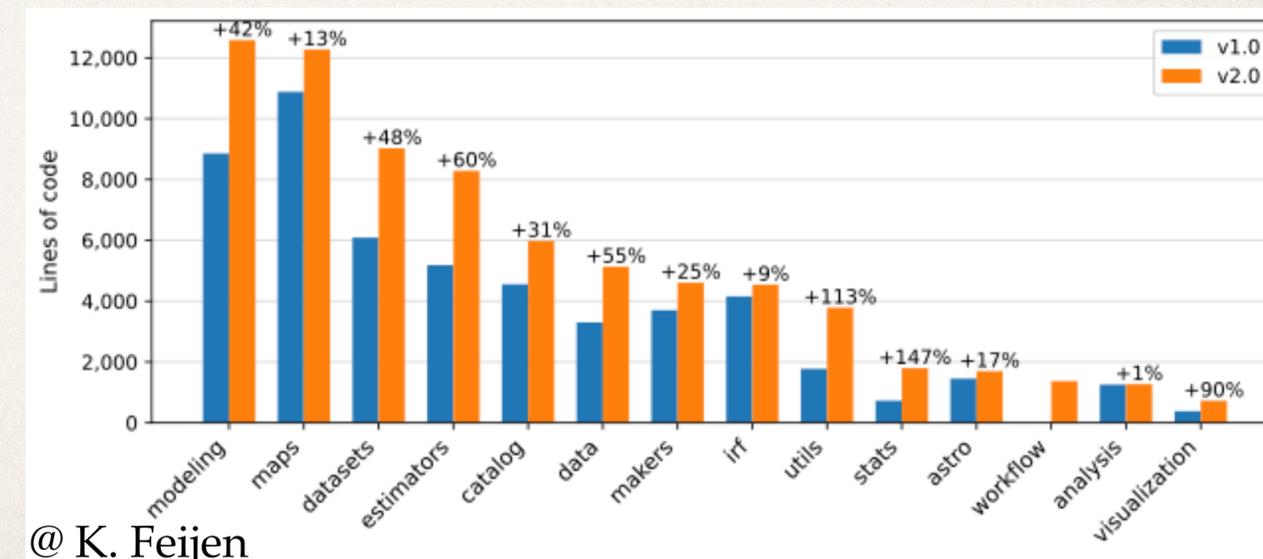
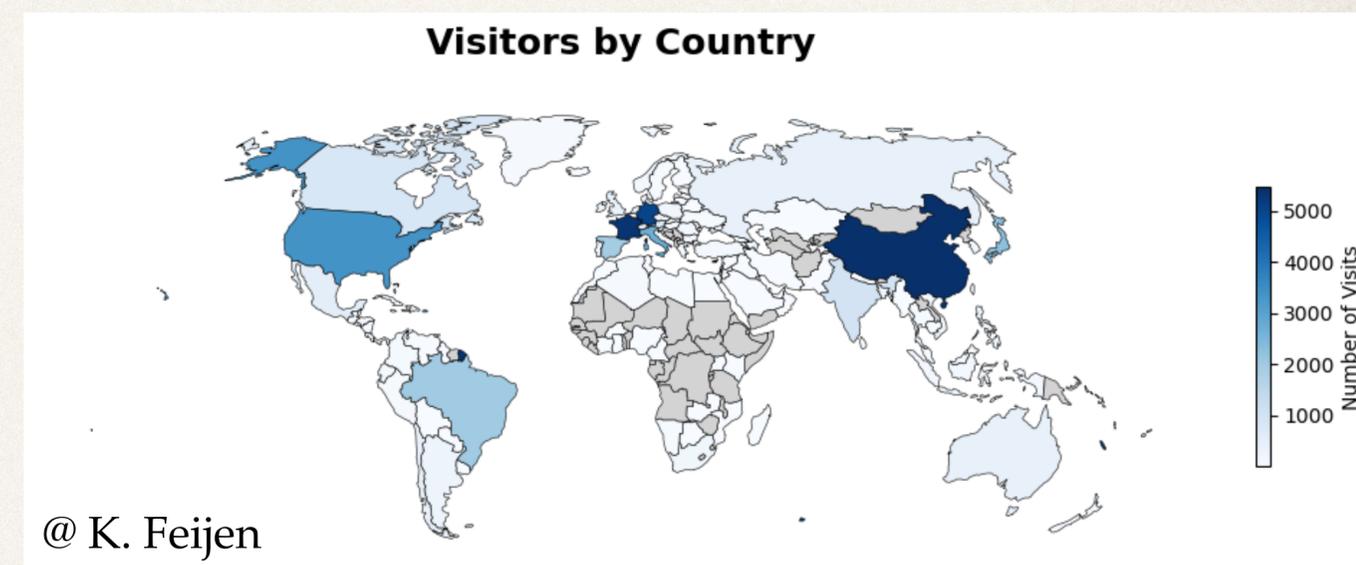
# Current status

## ❖ Releases:

- ❖ LTS 2.0: released 26 August, 2025
  - ❖ Bug fix 2.0.1: released 16th Dec, 2025
- ❖ Feature release 2.1: Expected March, 2025

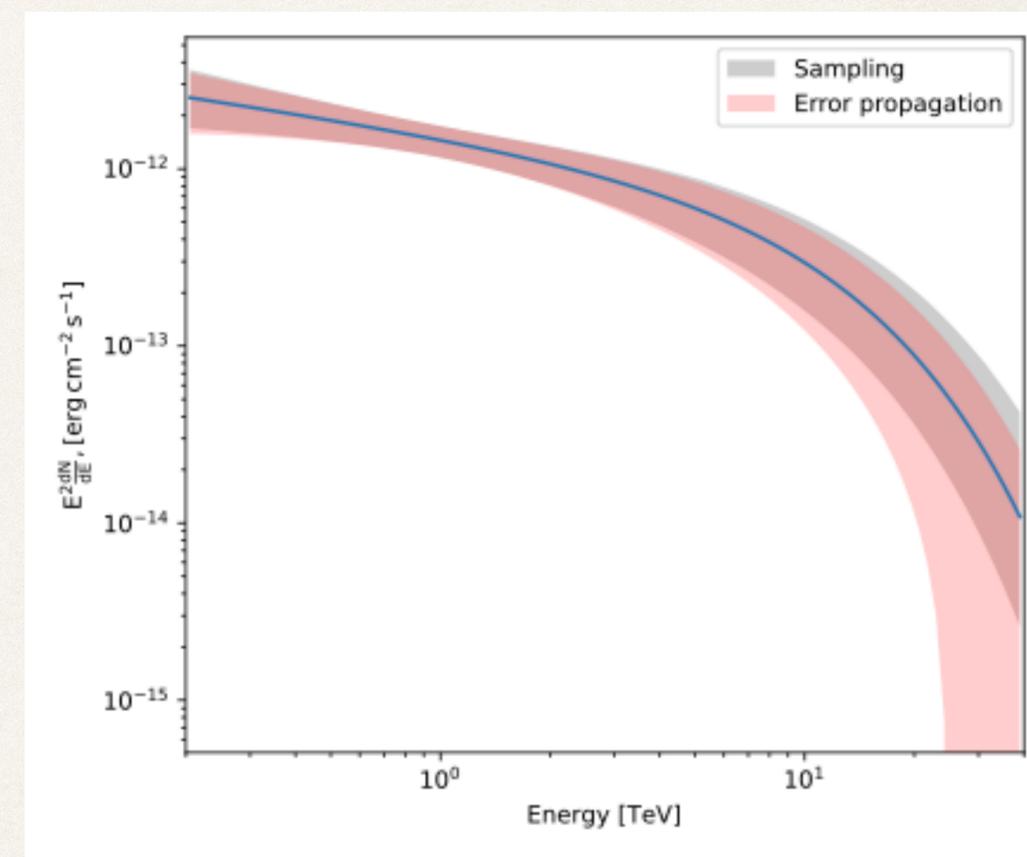
## ❖ Regular user calls:

- ❖ <https://github.com/gammapy/gammapy-meetings/tree/master/user-meetings>
- ❖ Get in touch with your usage!



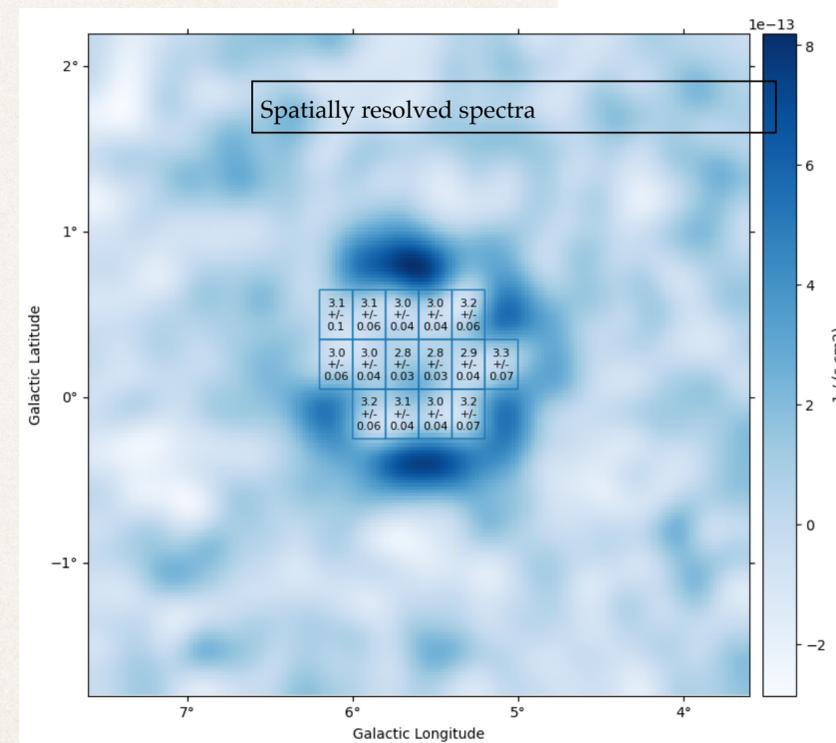
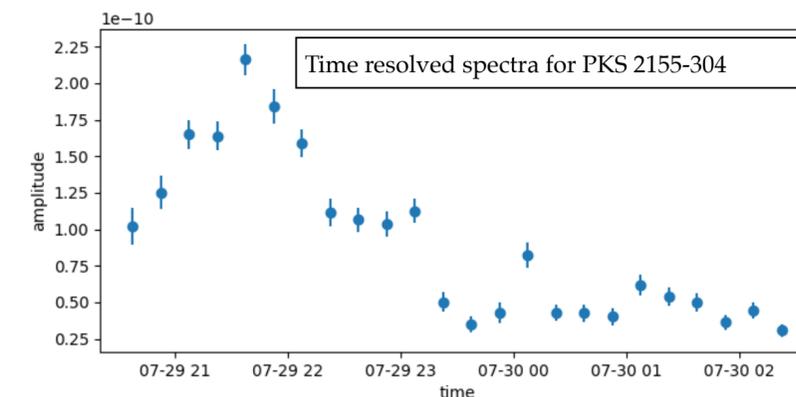
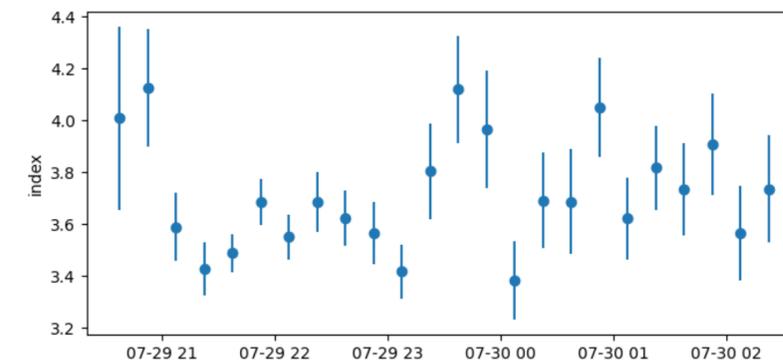
# Change in the butterfly plot

- ❖ Till version 1.3: Computation of flux errors performed with regular error propagation
  - Linear combination of multi-normal random variables
  - Problems far away from the decorrelation energy
- ➔ Version 2.x: Get uncertainty band by computing the envelope of flux for a Monte Carlo sample of the multi-normal variables
  - ➔ For user contributed models: Please try to vectorise your evaluations!



# Future plans

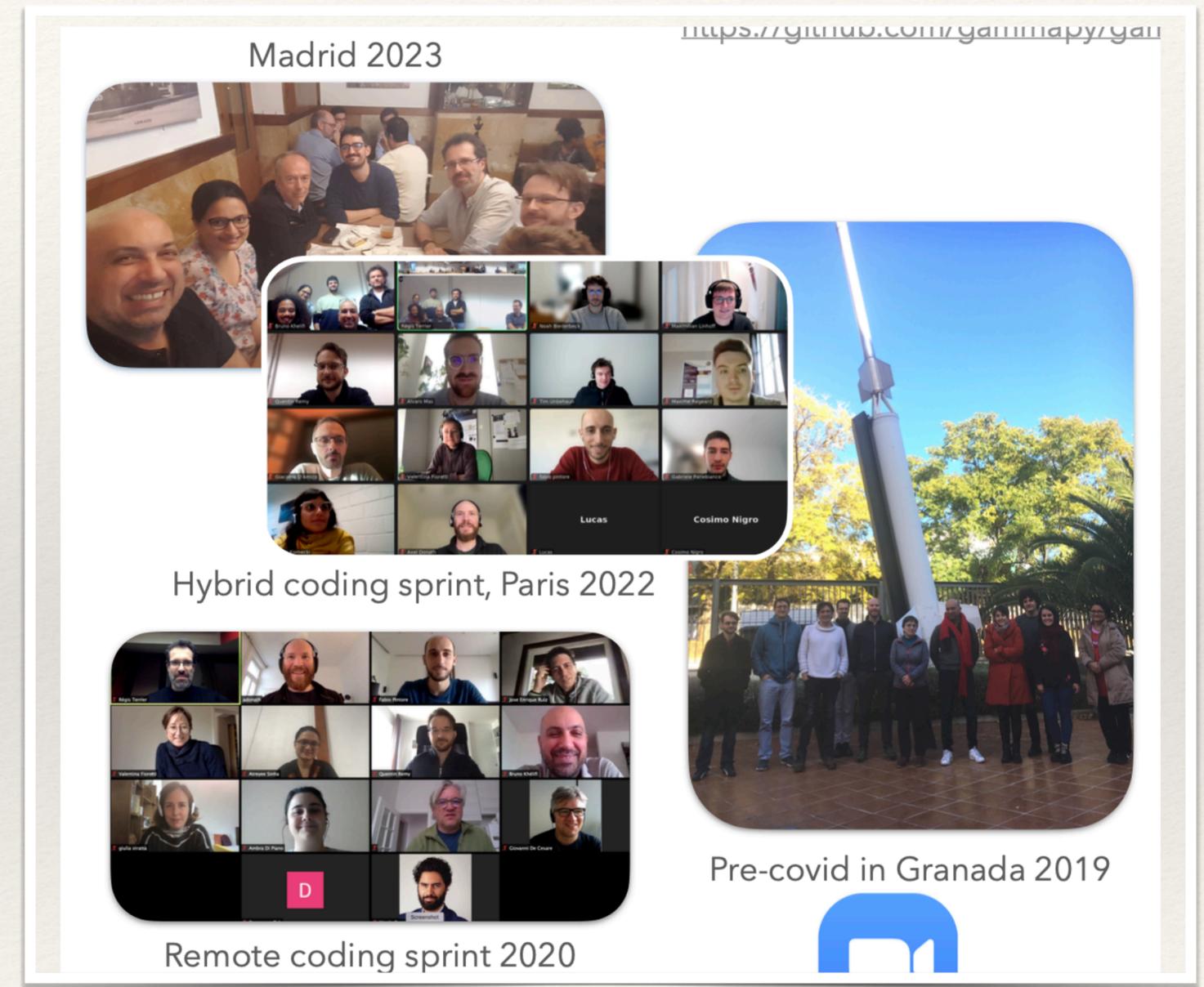
1. Improve Bayesian analysis framework
2. Support for unbinned analysis
3. Nuanced upper limit calculations
4. Support for unfolding
5. Time/spatially resolved spectroscopy
6. Improvement in `gammapy.astro` (dark matter, etc)
7. Handling systematic errors:
  - Partly supported with the priors



Comments, suggestions and contributions welcome!

# Join/Contact Us!

- ❖ Install Gammapy:
  - ❖ `conda install -c conda-forge gammapy`
  - ❖ New releases:
    - ❖ LTS: 2 yrs; major release: 6months; bug fixes: as required - released on Zenodo, PyPI, Conda
    - ❖ Latest dev version on GitHub
- ❖ Docs:
  - ❖ <https://docs.gammapy.org>
- ❖ Issues?
  - ❖ Slack: [gammapy.slack.com](https://gammapy.slack.com) (quick questions, immediate help)
  - ❖ GitHub issues: <https://github.com/gammapy/gammapy/issues> (feature requests & bug reports)



*Thank you!*